

Vlasov equation from Kaluza-Klein model and from the guiding center description of motion

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Introduction

- Newton law of gravitation $F = G \frac{m_1 m_2}{r^2}$
- Coulomb law: $F = \pm \frac{1}{4\pi\epsilon_0} \frac{e_1 e_2}{r^2}$
- ▷ In General Relativity bodies follow geodesics in space-time.
- ▷ The space-time is a manifold whose geometry depends on the distribution of particles and their energy by the Einstein's equation.
- ▷ There is not a direct interaction between particles.
- ▷ In the weak field approximation the effect on the space-time geometry acts as if a two body interaction applies.
- Even if the Coulomb law is so similar to the Newton gravitational law, it is believed that charges doesn't follow geodesics on space-time under electromagnetic forces.

Can charges follow geodesics, not on space-time, but on the phase-space manifold under electromagnetic and gravitational forces?

The guiding center description of motion

- The trajectory of a charge in e.m. field is seen to be helicoidal.
- Assume that exists a reference point in phase-space, called the guiding center, whose coordinates are X, U, t , from which the particle motion is seen to be closed and periodic (a loop in space as in the velocity space).
- If $\gamma \in S^1$ is the gyro-phase then the particle coordinates (x^α, u^α) in (a flat) phase space are expressed by

$$x^\alpha = X^\alpha + \rho^\alpha(\gamma), \quad u^\alpha = U^\alpha + \nu^\alpha(\gamma),$$

with $\alpha = 0, 1, 2, 3$.

- The guiding center coordinates are $(t, X, \gamma, \varepsilon, \mu)$. Explicitly: the time, t , the guiding center position, X , the gyro-phase, γ , the particle energy, ε , and the magnetic moment, μ .
- The guiding center transformation \mathcal{T} is the map from (τ, x, u) to $(t, X, \gamma, \varepsilon, \mu)$.
- The single particle lagrangian, L , is rewritten as

$$L = -p_\alpha u^\alpha = -P_\alpha U^\alpha - (m/e)\mu\gamma', \quad \text{with } P_\alpha = U_\alpha + (e/m)A_\alpha,$$

with $\alpha = 0, 1, 2, 3$ and summation convention.

- The equation of motion has to be obtained adopting the guiding center coordinates.

Hamiltonian formulation

▷ Define the e.m. canonical field by $(e/m)E_c = -\partial_t P - \nabla P_0$ and $(e/m)B_c = \nabla \times P$, and $b = B_c/|B_c|$.

▷ The Lagrange form is

$$\sigma = -d\varepsilon \partial_\varepsilon P \wedge dX - d\mu \partial_\mu P \wedge dX + dX \cdot B_c \times dX + (m/e)d\mu \wedge d\gamma$$

▷ with Hamiltonian form $\mathcal{H} = d\varepsilon \partial_\varepsilon P_0 + d\mu \partial_\mu P_0 - dX \cdot E_c$.

▷ the Poisson form is $\Pi = (e/m)\partial_\mu \wedge \partial_\gamma + \nabla \cdot (B_c/B_c^2) \times \nabla - \partial_\varepsilon \wedge b \cdot \nabla$

▷ equivalent to the Lie bracket

$$\{F, G\} = (e/m)(\partial_\mu F \partial_\gamma G - \partial_\gamma F \partial_\mu G) + B_c^{-2} \nabla F \cdot B_c \times \nabla G + (b \cdot \nabla F \partial_\varepsilon G - \partial_\varepsilon F b \cdot \nabla G)$$

▷ The obtained equations of motion are

$$\dot{X} = V_b b + \frac{E_c \times B_c}{B_c^2}, \quad \dot{\gamma} = -(e/m)\partial_\mu P_0, \quad \dot{\varepsilon} = V_b \frac{b \cdot E_c}{\partial_\varepsilon P_0}, \quad \dot{\mu} = 0$$

▷ If $\dot{\varepsilon} = 0$ and $\dot{X} = V$ then $E_c + V \times B_c = 0$ is the velocity law.

The lagrangian density on extended phase space

- The single particle lagrangian is an invariant (relativity principle) and should be extended to the phase-space coordinates if non canonical coordinates are adopted: if $z^a = (t, X, \gamma)$, $w_a = (P_0, -P, (m/e)\mu)$ and \tilde{g} is the metric tensor in the 5D extended phase-space, then

$$L = -w_a z'^a = -\tilde{g}_{ab} w^b z'^a,$$

with $a, b = 0, 1, 2, 3, 4$ and summation convention.

- The problem is that in the former eq. both \tilde{g} and w^b are unknown.
- The lagrangian density on extended phase-space is the following:

$$\ell a = f_m L - \frac{\tilde{\mathcal{R}}}{16\pi\tilde{G}},$$

being f_m the (single particle) distrib. of matter on the extended phase-space, $\tilde{\mathcal{R}}$ and \tilde{G} , the scalar curvature and the gravitational constant of the 5D extended phase-space, respectively..

- From ℓa , it'll be addressed
 - ▷ i) the motion of particles on extended phase-space (Lorentz and gravitational forces).
 - ▷ ii) the dynamics of fields (Maxwell's and Einstein's equations).
 - ▷ iii) the collisionless Boltzmann equation.

Einstein's equation on extended phase space

- The relativistic guiding center transformation in a flat phase-space is the translation of time, positions and velocities by quantities that depend on $\gamma \in S^1$:

$$x^\alpha = X^\alpha + \rho^\alpha(\gamma) \quad (1)$$

$$u = U + \nu(\gamma). \quad (2)$$

- The misleading symmetry is the condition: $(m/e)\mu\gamma' = 1 - U^\alpha U_\alpha$.
(it corresponds to $(m/e)\mu\dot{\gamma} = \varepsilon - \mathcal{E}$ in non relativistic regimes)
- The relativistic Lagrangian has the same form after the GC transformation that is a translation:

$$L = -1 - (e/m)A_\alpha(x^\beta)u^\alpha = -1 - (e/m)A_\alpha(X^\beta)U^\alpha,$$

with $\alpha = 0, 1, 2, 3$ and summation convention.

- The form of the Lagrangian remains the same after a translation on phase-space \Rightarrow The extended energy momentum tensor, \tilde{T}_{ab} , is the conserved Noether current.
- From the 5d **Hilbert Einstein term** in the lagrangian density, $\ell a = f_m L - \tilde{R}/(16\pi\tilde{G})$, and from varying the metric tensor,:

$$\tilde{G}_{ab} = 8\pi\tilde{G} \tilde{T}_{ab}. \quad \text{Einstein's equation on the 5D extended phase space} \quad (3)$$

Kaluza Klein metric and unification of forces

- In 5D: $L = -\tilde{g}_{ab}w^a z'^b$, $a, b = 0, 1, 2, 3, 4$ and $\ell a = f_m L - \tilde{R}/(16\pi\tilde{G})$.
- The misleading symmetry: $U^a U_a = 1$.
- The following KK metric tensor is adopted:

$$\tilde{g}_{ab} = \begin{vmatrix} g_{\alpha\beta} - k_G^2 A_\alpha A_\beta & k_G^2 (m/e)^2 \mu A_\alpha \\ k_G^2 (m/e)^2 \mu A_\beta & -k_G^2 (m/e)^4 \mu^2 \end{vmatrix}. \quad (4)$$

being $k_G = \sqrt{16\pi G}$.

- being $\sqrt{|\tilde{g}|} = \sqrt{-g}(m/e)^2 k_G \mu$ and $\tilde{G} = G \int (m/e)^2 k_G \mu \tilde{J}_{\mathcal{P}} d\gamma d\varepsilon d\mu$.
- thus,

$$S_{\text{field}} = - \int \frac{\tilde{\mathcal{R}}}{16\pi\tilde{G}} \sqrt{|\tilde{g}|} d^7 z = - \int \sqrt{-g} dt d^3 X \frac{R}{16\pi G} - \int \sqrt{-g} dt d^3 X \frac{F_{\alpha\beta} F^{\alpha\beta}}{4}. \quad (5)$$

- The terms in the lagrangian density (on the space-time) are the desired ones: e.m. and gravitational, but they are referring to fields on (t, \mathbf{X}) where \mathbf{X} is the guiding center position and it doesn't indicate the position of a particle \Rightarrow the present theory is NON LOCAL

Boltzmann's and Vlasov's equations without BBGKY

- By writing $L = \mathcal{S}'$, as the proper time derivative of the principal Hamilton function, \mathcal{S} , then it is possible to recognize a Legendre's transformation in

$$\ell a = f_m \mathcal{S}' - \frac{\tilde{\mathcal{R}}(f_m, \mathcal{S})}{16\pi\tilde{G}}. \quad \text{like for} \quad L(q, \dot{q}) = p\dot{q} - H(p, q) \quad (6)$$

- The distribution of matter, f_m , is the conjugate momentum of the principal Hamilton function.
- By varying such lagrangian density on 5D extended phase-space with respect to \mathcal{S} and \mathcal{S}' , then the Euler-Lagrangian equation is

$$f'_m = -\frac{\partial_{\mathcal{S}} \tilde{\mathcal{R}}}{16\pi\tilde{G}}. \quad (7)$$

- The former is a Boltzmann-like equation, describing the evolution of the distribution function, with a *RHS* that depends on the geometry instead of collisions.
- If $\tilde{\mathcal{R}}$ doesn't depend on \mathcal{S} then the Vlasov equation is obtained (describing the equilibrium).
- BBGKY is overcome because there aren't collisions but geodesics, the two body interaction could be the effect of a weak field approximation that may not be applicable in a strong magnetized plasma. In such case the fluid picture is obtained without the need of truncations.

Conclusions

- ▷ The density lagrangian on the extended phase-space has been proposed:

$$\ell a = f_m L - \frac{\tilde{\mathcal{R}}}{16\pi\tilde{G}},$$

for deriving

- Lorentz's force law by varying the single particle lagrangian with respect positions and velocities.
 - Maxwell's and Einstein's equations by varying it with respect to the metric tensor on the extended phase space.
 - (Collisionless) Boltzmann's equation by varying it with respect to the principal Hamilton function and the single particle Lagrangian.
- ▷ Particles follow geodesics on extended phase-space (unification of forces).
 - ▷ The single particle distribution is the conjugate momentum of the principal Hamilton function.
 - ▷ BBGKY is overcome.
 - ▷ At the equilibrium the Vlasov's equation is obtained.

Thanks a lot

References:

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- Valvo L. and Di Troia C., 46th EPS Conference on Plasma Physics (2019) contribution, "Hamiltonian Formulation of the Non-perturbative Guiding Centre Equation"
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Backup slides

- The lagrangian for describing electrodynamics (without gravitation) for plasmas is the single particle lagrangian, $L(t, x, v)$, times the distribution function of charges, $f_m(t, x, v)$, added to the e.m. lagrangian. The action is often expressed like:

$$S_{\text{plasma}} = \int dt d^3x d^3v f(t, x, v) L(t, x, v) - \int \frac{F_{\alpha\beta} F^{\alpha\beta}}{4} dt d^3x, \quad (8)$$

- There is an unbalance in the action (8) between the matter action and the field one.
- In principle, for restoring the symmetry between the two lagrangians, it should be simple to think at an action written as

$$S_{\text{plasma}} = \int dt d^3x d^3v \mathcal{L}_{\text{plasma}}, \quad (9)$$

where $\mathcal{L}_{\text{plasma}} = f_m(t, x, v) L(t, x, v) + \text{"somethingnew"}$ and the property that

$$\int \text{"somethingnew"} d^3v = -\frac{F_{\alpha\beta} F^{\alpha\beta}}{4}. \quad (10)$$

Two key-points: General relativity and BBGKY hierarchy

▷ General relativity in 4 steps

- Translational symmetry: if $x^\alpha = X^\alpha + \rho^\alpha$ (flat space time) then the lagrangian density remains the same.
- Noether current: the energy momentum tensor $T^{\alpha\beta}$ is conserved due to the translational symmetry.
- The density lagrangian density describing the gravitational field is the Hilbert-Einstein one

$$\mathcal{L}_{\text{HE}} = -\frac{R}{16\pi G}$$

- By varying the metric tensor, the Einstein's equation is obtained:

$$G_{\alpha\beta} = 8\pi G T_{\alpha\beta}. \quad (11)$$

▷ BBGKY for N interacting particles in 2 steps

- If the Hamiltonian is $\mathcal{H} = \sum_{i=1}^N p_i^2/(2m) + \sum_{i=1}^N U(r_i) + \sum_{i<j}^N V(|r_i - r_j|)$,
- then, the single particle distr. func. depends on the 2-particles distr. func. and so on and so forth until the (N-1)-particles distr. func. depends on the N-particles distr. func.
- that means that you cannot know the single-particle distr. func. if you don't already know the N-particles distr. func. For such reason the transport equations are a truncated set of equations, and the fluid picture is an approximation of the kinetic one.

Lorentz's force law solutions of motion

- A charge e with mass m that (non-relativistically) moves in a externally given and static e.m. field is classically described by the system of equations:

$$\dot{x} = v; \quad (12)$$

$$\dot{v} = (e/m)(E + v \times B) \quad (13)$$

within **eulerian description** $\dot{x} = v(t, x, \alpha)$ with $\dot{\alpha} \cdot \nabla_{\alpha} v = 0$. Substituting $E = -\partial_t A - \nabla \Phi$ and $B = \nabla \times A$, eq. (13) becomes

$$\partial_t v + \dot{x} \cdot \nabla v + \dot{\alpha} \cdot \nabla_{\alpha} v = (e/m)(-\partial_t A - \nabla \Phi + v \times \nabla \times A). \quad (14)$$

or

$$\partial_t [v + (e/m)A] + \nabla [v^2/2 + (e/m)\Phi] = v \times \nabla \times [v + (e/m)A]. \quad (15)$$

Equivalent to **the velocity law**:

$$E_c + v \times B_c = 0, \quad (16)$$

if the **canonical e.m. fields** are defined as

$$(e/m)E_c = -\partial_t p - \nabla \varepsilon \quad \text{and} \quad (e/m)B_c = \nabla \times p \quad (17)$$

The guiding center description of motion

- The guiding center description of motion is not eulerian nor lagrangian.

$$x = X + \rho(\gamma), \quad (18)$$

$$\dot{x} = V + \sigma(\gamma). \quad (19)$$

- $\dot{X} = V(t, X, \mu, \varepsilon)$ with $E_c + V \times B_c = 0$.
- If $\mathcal{E} = V^2/2 + (e/m)\Phi(t, X)$, then the cyclotron frequency, $\dot{\gamma}$, is chosen to be

$$(m/e)\mu\dot{\gamma} = \varepsilon - \mathcal{E} \quad (20)$$

- the guiding center momentum is $P = V + (e/m)A(t, X)$.
- The lagrangian doesn't explicitly depend on the gyro-phase, γ is cyclic.
- The constant of motion for such symmetry is the magnetic moment, μ .

The guiding center Lagrangian

- The single (charged) particle Lagrangian, L , must be gauge independent: if

$$L \rightarrow L + \dot{\mathcal{S}}, \quad (21)$$

the equations of motion are the same.

- L must be a scalar: if the index A runs over the extended phase-space dimensions then

$$L = \gamma_A \dot{z}^A \quad (22)$$

▷ Canonical coordinates:

$$z^A = (t, x, v) \quad \text{and} \quad \gamma_A = (-\varepsilon, p, 0).$$

▷ Non canonical, guiding center coordinates:

$$Z^A = (t, X, \gamma, \mu, \varepsilon) \quad \text{and} \quad \Gamma_A = (-\mathcal{E}, P, -(m/e)\mu, 0, 0).$$

- $L = p \cdot \dot{x} - \varepsilon = P \cdot \dot{X} - \mathcal{E} - (m/e)\mu\dot{\gamma} \iff p \cdot \dot{x} = P \cdot \dot{X} \text{ if } \varepsilon = \mathcal{E} + (m/e)\mu\dot{\gamma}.$
- If $\mathcal{S} = (m/e)\mu\gamma$ then guiding center Lagrangian is

$$L_{gc} = L + \dot{\mathcal{S}} = P \cdot \dot{X} - \varepsilon + (m/e)\mu\dot{\gamma}.$$